

# Evidence for Rapid Photodestruction of Interstellar Dust by Radiation from Gamma-Ray Burst 120119A

*Daniel A. Perley (Caltech), Adam N. Morgan (UC Berkeley), S. B. Cenko (GSFC), J. S. Bloom (UC Berkeley), A. Cucchiara (GSFC), J. W. Richards (UC Berkeley), A. V. Filippenko (UC Berkeley), J. B. Haislip (UNC Chapel Hill), A. LaCluyze (UNC Chapel Hill), A. Corsi (GWU), A. Melandri (INAF; Liverpool JMU), B. E. Cobb (GWU), A. Gomboc (U. of Ljubljana), A. Horesh (Caltech), B. James (UC Berkeley); W. Li (UC Berkeley), C. G. Mundell (Liverpool JMU), D. E. Reichart (UNC Chapel Hill), I. Steele (Liverpool JMU)*

We present a time-resolved analysis of the dust properties along the line of sight to gamma-ray burst (GRB) 120119A at redshift  $z = 1.73$ . Late-time photometric observations of the burst's optical afterglow show evidence for a moderate column of dust ( $A_V \sim 1.1$  mag) with an extinction curve similar to, but statistically distinct from, dust seen along Small Magellanic Cloud sightlines. However, our earliest observations exhibit a significant red-to-blue color change in the first  $\sim 200$  s after the burst at levels heretofore unseen in GRB afterglows. This color change, which is coincident with the final phases of the prompt emission, is a hallmark prediction of the photodestruction of dust in GRB afterglows. We test whether dust-destruction signatures are significantly distinct from other sources of color change, namely a change in the intrinsic spectral properties of the GRB itself. We find that a time-varying spectrum within the constraints of accepted GRB models cannot alone adequately describe the observed color change, and allowing for dust destruction (via a time-varying  $A_V$ ) significantly improves the fit. While not definitively ruling out other possibilities, this event provides the best support yet for the direct detection of dust destruction in the local environment of a GRB.