

# The Mid-Infrared Camera and Spectrometer for SPICA: General overview and Taiwan's contribution

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# Outline

- The MCS consortium & Instrument Specifications
- Highlights of the three science cases
  - The Drama of Galaxy Formation
  - The Recipe for Planet Formation
  - The Transmigration of Dust in the Universe
- Taiwan's contribution to MCS
  - Scientific underpinning and definition of instrument specs
  - Filters
  - Calibration lamp
  - Detectors

# The MCS consortium & Instrument Specifications

## MCS Consortium:

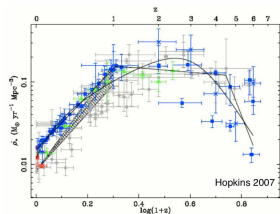
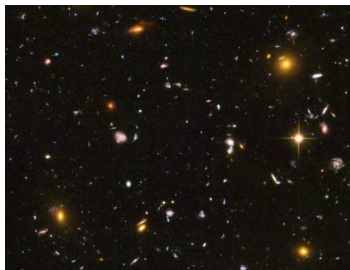
- Japan & Taiwan
- PI: H. Kataza
- Project Scientist: T. Onaka
- > 90 Co-Is

## Instrument Specifications:

- WFC
  - FOV:  $5' \times 4'53''$
  - $20 - 38 \mu\text{m}$  (WFC-L);  $5 - 25 \mu\text{m}$  (WFC-S)
  - $R = 10$  (WFC-L);  $R = 5$  (WFC-S)
  - filters not decided yet, may include low ( $R = 50$ ) and medium ( $R = 200$ ) resolution grisms & narrow-band filters
- HRS-L;  $12 - 18 \mu\text{m}$ ;  $R = 20,000 - 30,000$
- MRS-S,-L;  $12.2 - 37.5 \mu\text{m}$ ;  $R = 1000 - 2000$

# The Drama of Galaxy Formation (I)

- Star formation over cosmic time
- Role of AGN in galaxy evolution
- Assembly of Milky-Way type galaxies

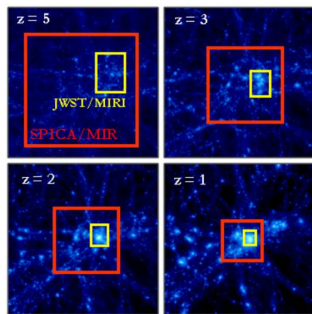


- For  $4 < z < 10$  optical/NIR tracers of SF fall in MCS range, e.g.  $H\alpha$ ,  $Ly\alpha$
- Does SFR really peak at  $z \approx 2 - 3$ ?



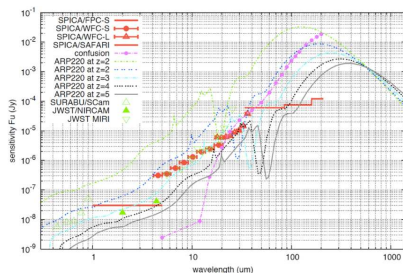
# The Drama of Galaxy Formation (III)

Galaxy environments:



- superior mapping compared to JWST/MIRI
  - FOV
  - sensitivity  $> 20 \mu\text{m}$

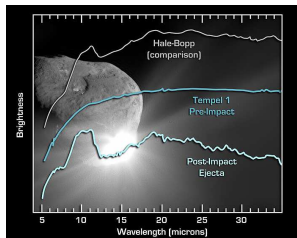
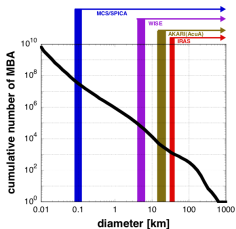
Spectral energy distribution:  
SPICA MCS and SAFARI  
synergy



- Slitless spectroscopy

# The Recipe for Planet Formation (I)

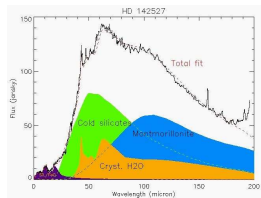
## Solar system objects: surveys and mineralogy



- IRAS: 2470 Main Belt Asteroids
- WISE: 130,000 MBAs
- SPICA/MCS:  $10^7-8$  MBAs?

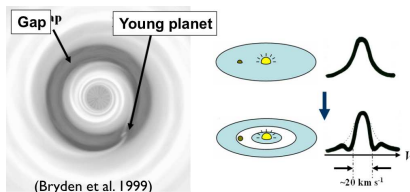
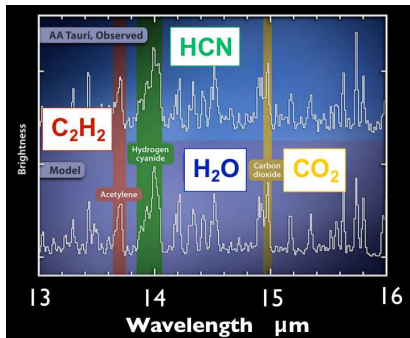
⇒ Size distribution

## Combination with SAFARI:



# The Recipe for Planet Formation (II)

## Gas dissipation in planet forming disks

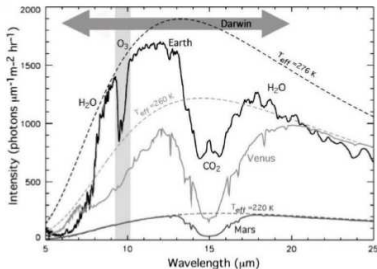
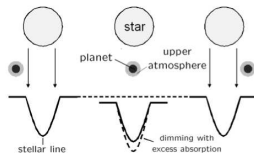




# The Recipe for Planet Formation (III)

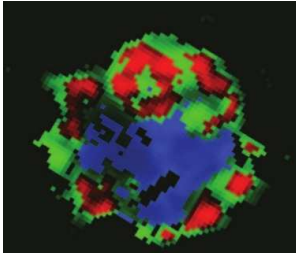
## Atmospheres of transiting exoplanets

- MIR coronagraphy is desired for observing exoplanets with ages close to our Solar System
  - Difficult to observe from the ground
- A combination with spectroscopy will allow for studies of
  - atmospheric composition
  - climate
  - exoplanet diversity
- In connection with SCI team

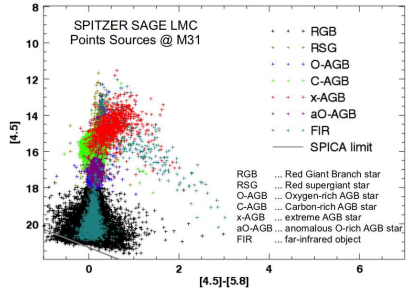
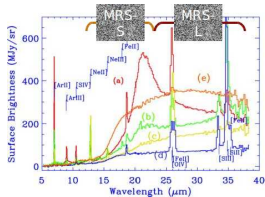


# The Transmigration of Dust in the Universe (I)

Sources of dust on a galactic scale: AGB stars and SNe/SNR



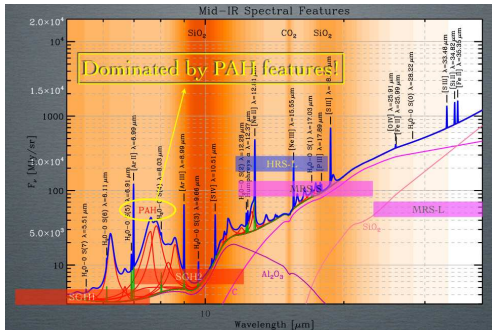
Cas A (Rho et al. 2008)



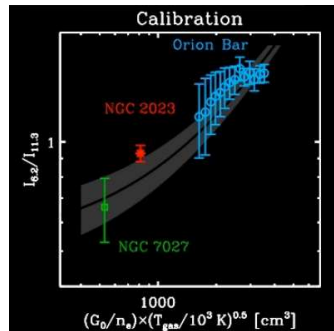
Dust production by AGB stars  
in M31

# The Transmigration of Dust in the Universe (II)

## The properties of dust and gas in the ISM

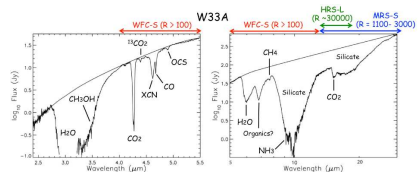


- PAH bands
- molecular lines
- atomic lines
- solid state features



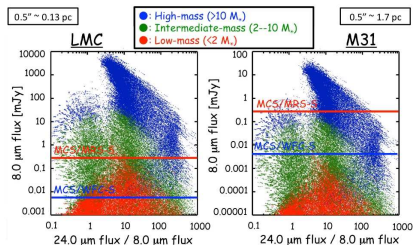
# The Transmigration of Dust in the Universe (III)

## Circumstellar material around YSOs



## Spectroscopic detection of extragalactic YSOs with MCS

- 1172 YSO candidates in the LMC (Gruendl & Chu, 2009): 277 observed with IRS
- 282 YSO candidates in the SMC
- M31 ?



# Taiwan's contribution to MCS

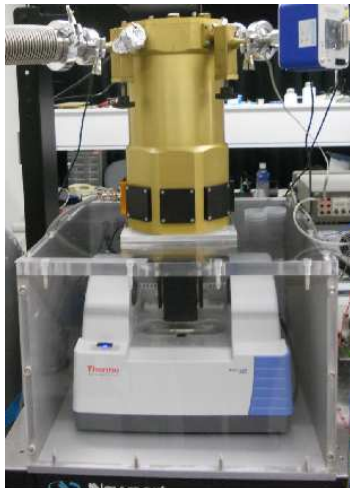
## Taiwanese IR astronomy community:

- Strengthening science case
- Further instrument specifications (e.g. filter specification)
- Representative observations

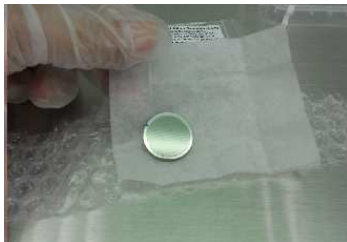
## OIR lab ASIAA:

- Broadband filter development
- Calibration microlamp development
- Detector module development
  - Procurement and testing of  $2k \times 2k$  Si:As arrays and  $1k \times 1k$  Si:Sb arrays
  - Assembly of flight modules

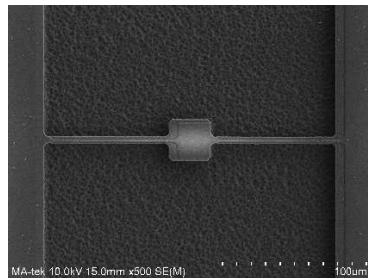
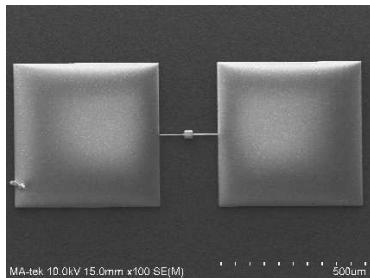
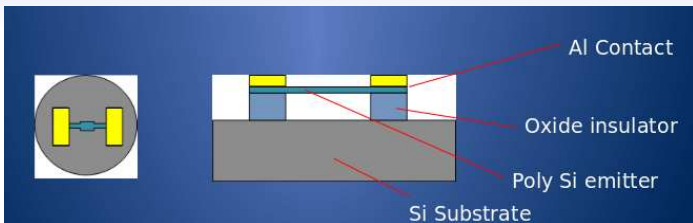
# Broadband filter development



- Specification of filters needed
- Test setup



# Calibration microlamp development



# Summary

MCS is being developed by a Japanese/Taiwanese consortium

MCS will contribute to three main science cases:

- The Drama of Galaxy Formation
- The Recipe for Planet Formation
- The Transmigration of Dust in the Universe

Taiwan's contribution to MCS:

- Scientific underpinning and definition of instrument specs
- Filters
- Calibration lamp
- Detectors