



Radiative Feedback of Massive Stars

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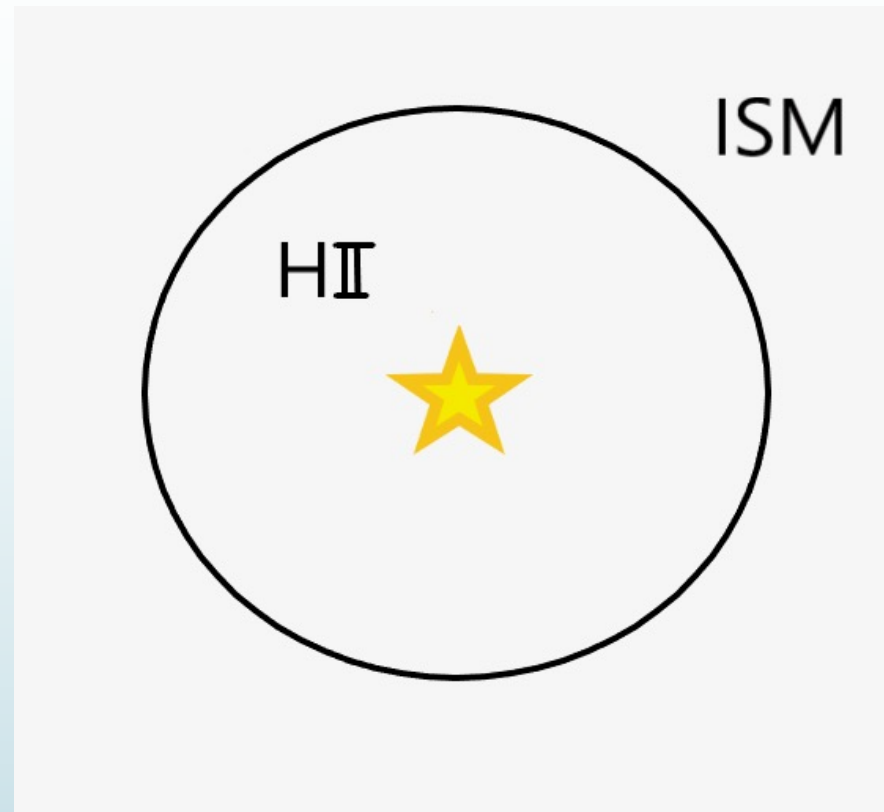
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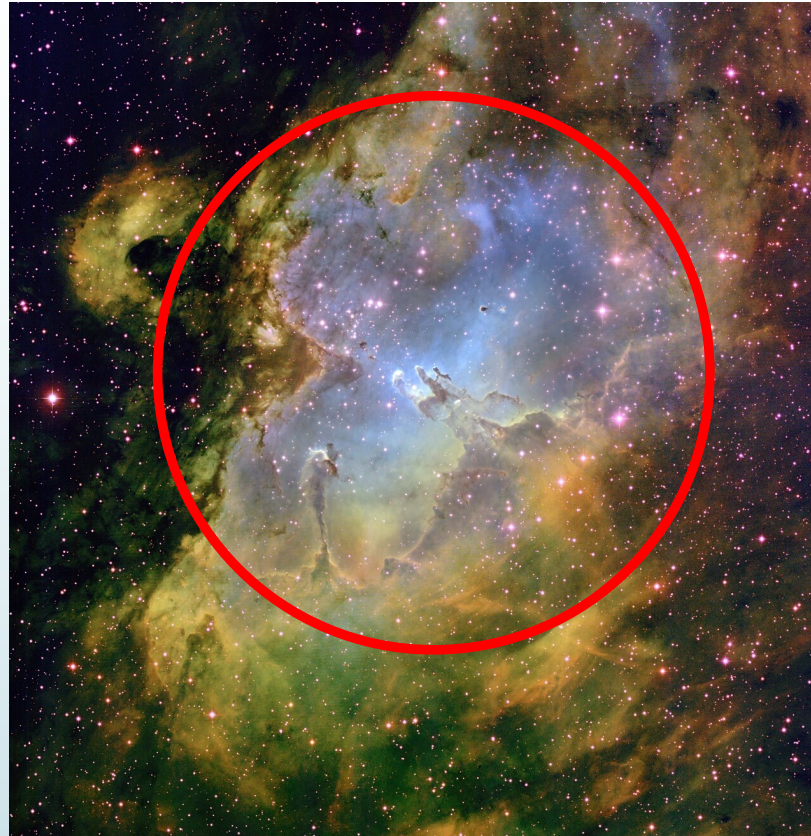
Outline

- ▶ Introduction
- ▶ Method and result
- ▶ Conclusion and Future work

Introduction



Eagle Nebula



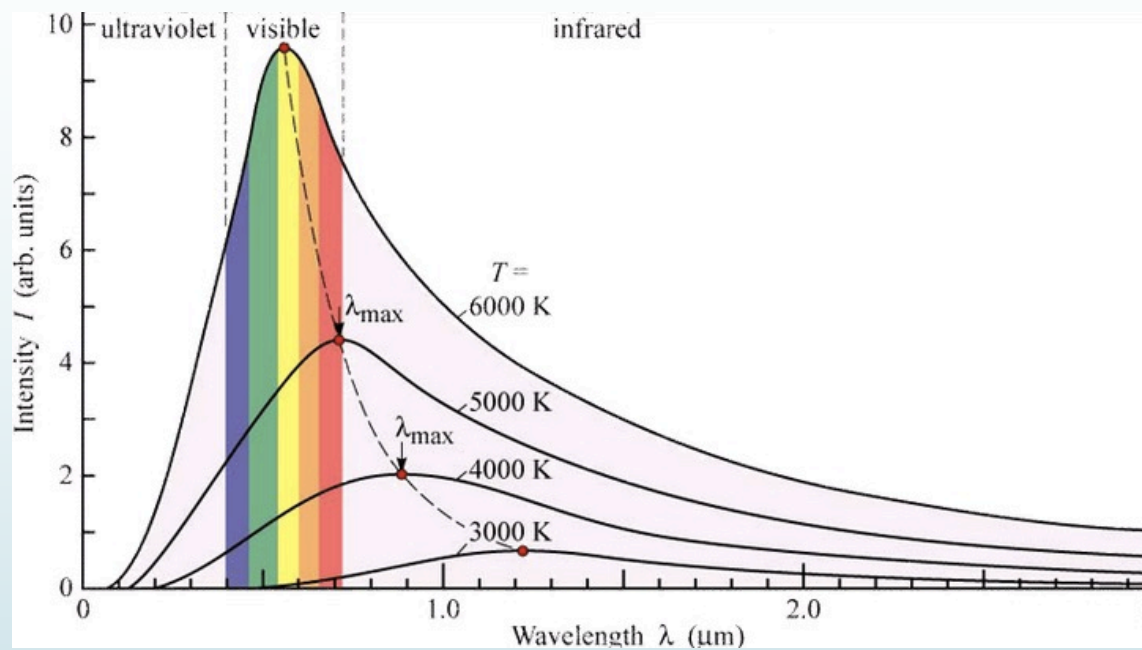
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Why?

- ▶ The ionization energy is 13.6eV
- ▶ Equal to 136,000k

Black Body Radiation





Photoionization Rate

$$n_{\xi} \int \frac{L_{\nu}}{4\pi r^2} \frac{a_{\nu}}{h\nu} d\nu$$

(Number of neutral H) X (number of photons from source)

Recombination Rate

$$n_e \cdot n_p \cdot \alpha$$



Equilibrium State

$$n_H \xi \int \frac{L_\nu}{4\pi r^2} \frac{a_\nu}{h\nu} d\nu = \text{Photoionization Rate}$$

$$n_e \cdot n_p \cdot \alpha = \text{Recombination Rate}$$

$$n_H \xi \int \frac{L_\nu}{4\pi r^2} \frac{a_\nu}{h\nu} d\nu = n_e \cdot n_p \cdot \alpha$$




Simulation Tool

RAMSES

- ▶ The AMR service routines
- ▶ the particle mesh routines
- ▶ the poisson solver
- ▶ the hydrodynamical routines

PymSES

an analysis library written in Python for RAMSES
outputs: <http://irfu.cea.fr/Projets/PYMSES/>

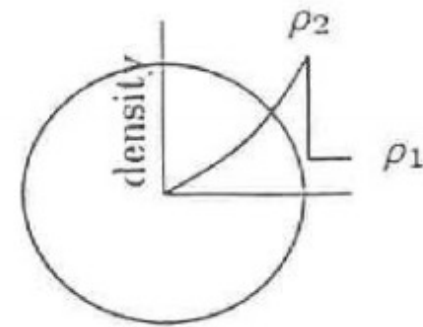


Test Problem:
Sedov2d



Sedov Blastwave

blast wave
(energy conserving)



Dimensional Analysis

- Amount all relevant parameters we pick 3 of them as scaling parameters

Parameters: E, t, ρ

Dimensions: $\left\{\frac{ML^2}{T^2}\right\}, \{T\}, \left\{\frac{M}{L^3}\right\}$

Fundamental dimension: $\{L\}, \{T\}, \{M\}$

- nondimensionalize the radius to form dimensionless groups

$$\xi = r \rho^k t^h E^n$$

Length dimension: $L = k - 3h + 2n = 0$

Mass dimension: $M = h + n = 0$

Time dimension: $T = k - 2n = 0$



Dimensional Analysis

$$\xi = r \left(\frac{\rho}{Et^2} \right)^{1/5}$$

$$r = \xi \left(\frac{Et^2}{\rho} \right)^{1/5}$$

$$\frac{dr}{dt} = v = \frac{2}{5} \xi \left(\frac{E}{\rho t^3} \right)^{1/5}$$



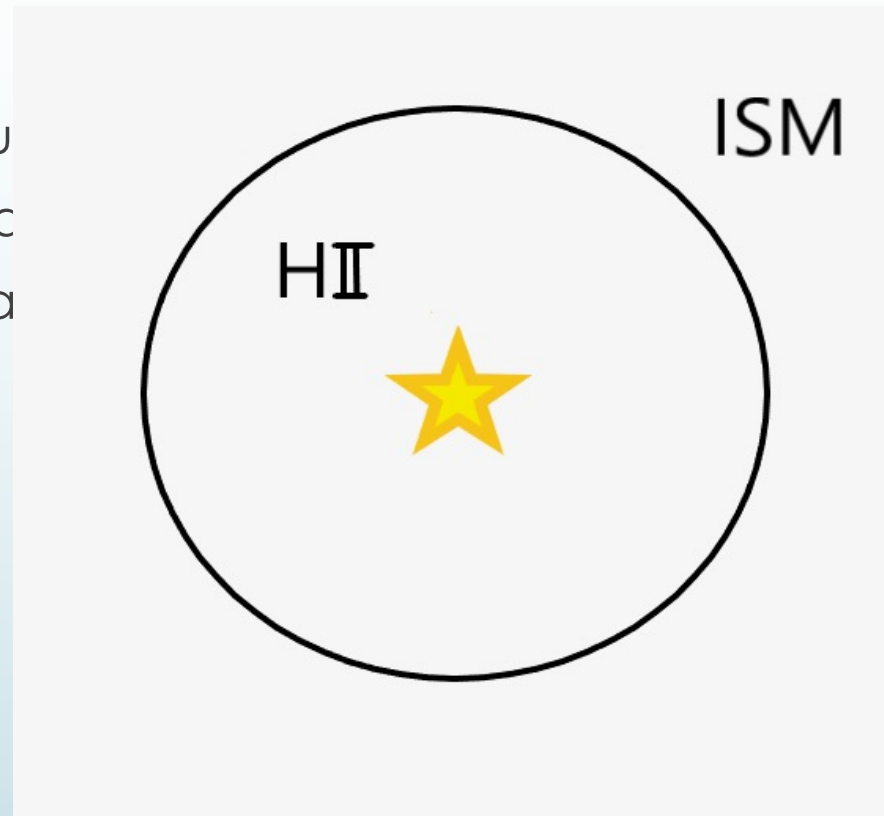
Preliminary

- ▶ Physics of HII region
- ▶ Ramses code
- ▶ Self-similar solution of Sedov blastwave

This is just a beginning !

Future Work

- ▶ Pu
- ▶ Ac
- ▶ Pa





Thank you for your listening ~